

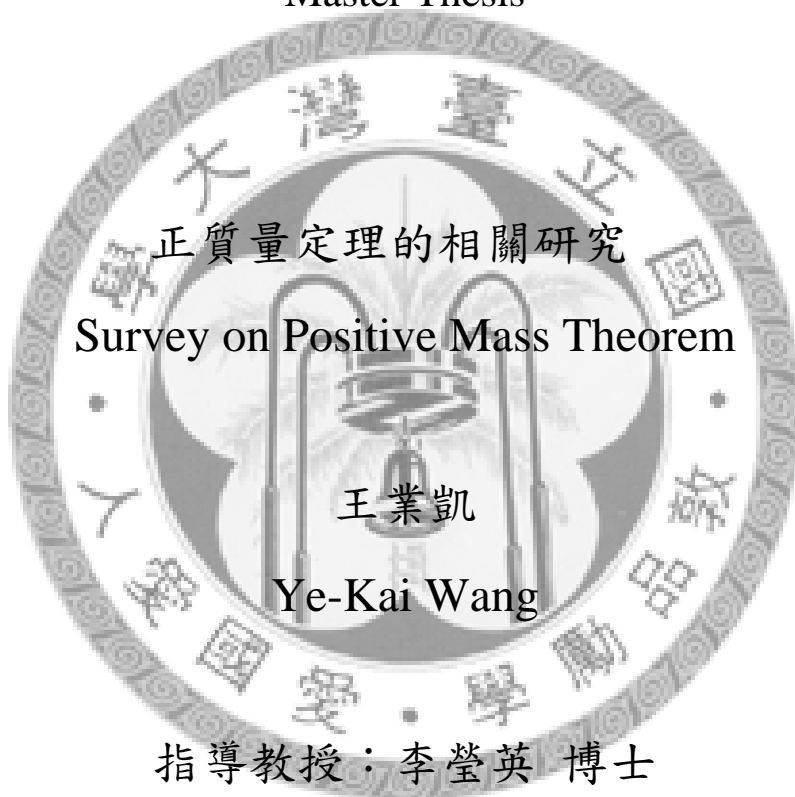
國立臺灣大學數學研究所

碩士論文

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正質量定理的相關研究  
Survey on Positive Mass Theorem

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## 摘要

這篇文章介紹 quasilocal mass 的相關研究，主要是 Brown-York and Liu-Yau quasilocal mass 的定義以及如何證明它們是非負的。第一個重要發現是史宇光與譚聯輝證明了 time-symmetric 的情形，接著劉秋菊與丘成桐給出一般情形下的證明，最後我們討論王慕道與丘成桐最近提出的 quasilocal mass 的定義。

關鍵字: 正質量定理



ABSTRACT. This note surveys the definition of quasilocal mass and its positivity. In particular, we focus on Brown-York and Liu-Yau quasilocal mass. We first present Shi and Tam's result on the positivity of quasilocal mass in the Riemannian case and then Liu and Yau's approach to the general case. Finally, we mention a modification of Liu-Yau quasilocal mass by Wang and Yau.

Key word: quasilocal mass, positive mass theorem, quasi-spherical, isometric embedding, Jang's equation, static mean curvature



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# 1 Introduction

In this note, we survey the definition of quasilocal mass and its positivity.

We first review the history of the positive mass theorem. In Einstein's theory of general relativity, the spacetime is a 4-dimension Lorentzian manifold  $(N, g)$  satisfying Einstein equation  $R_{\alpha\beta} - \frac{1}{2}Rg_{\alpha\beta} = 8\pi G T_{\alpha\beta}$ , where  $G$  is Newtonian constant and  $T_{\alpha\beta}$  is the (symmetric)energy-momentum tensor. Since general relativity can be viewed as an extension of classical Newtonian mechanics, it is desirable to define the notion of mass, energy, momentum, and angular momentum.

There are, however, several fundamental difficulties. First, the underlying manifold is unknown. All physical observations up to now are only local measurements compared with the scale of the universe and give no information about the topology of the universe. Second, Einstein equation is a nonlinear hyperbolic system of 10 degrees of freedom. The knowledge and techniques for such systems are limited. Third, there is no precise definition on how to relate the distribution of matter to  $T_{\alpha\beta}$ . It seems impossible to treat the general case directly.

A natural approach is to start from special cases. One case that has the longest history and is most extensively studied is the isolated gravitating system. Its origin could be traced back to Schwarzschild's model of the gravitational field of a single star in 1916. Mathematically, the isolated gravitating system is represented by an asymptotically flat spacelike hypersurface in the spacetime.

**Definition.** A 3-dimensional manifold  $M \subset (N, g)$  is asymptotically flat if for some compact set  $C$ ,  $M \setminus C = \cup_{i=1}^p M_i$  such that each  $M_i$  is diffeomorphic to  $\mathbb{R}^3 \setminus B_0(R_i)$ . Under this diffeomorphism, the metric is required to be of the form

$$g_{ij} = \delta_{ij} + a_{ij},$$

where  $a_{ij} = O(r^{-1})$ ,  $\partial_k a_{ij} = O(r^{-2})$ ,  $\partial_k \partial_l a_{ij} = O(r^{-3})$ . Moreover, the second fundamental form  $p_{ij}$  of  $M$  decay as  $p_{ij} = O(r^{-2})$ ,  $\partial_k p_{ij} = O(r^{-3})$ . The triple  $(M, g_{ij}, p_{ij})$  is called an *initial data set*.

One usually requires that the energy-momentum tensor satisfies certain energy conditions. We say that  $T_{\alpha\beta}$  satisfies the dominant energy condition if for any orthonormal frame  $\{e_\alpha | \alpha=0, 1, 2, 3\}$  at  $p \in M$ , with  $e_0$  normal to  $M$ ,

$$T_{00}^2 \geq \left( \sum_{i=1}^3 T_{0i}^2 \right), \quad (\text{T1})$$

and

$$T_{00} \geq |T_{\alpha\beta}|. \quad (\text{T2})$$

If only (T1) holds, we say  $T_{\alpha\beta}$  satisfies the weak energy condition.

In 1962, Arnowitt, Deser, and Misner defined the total energy and total momentum of an asymptotically flat manifold. They are defined on each asymptotically end  $M_l$

$$E_l = \frac{1}{16\pi G} \lim_{R \rightarrow \infty} \int_{S_R} (g_{ij,j} - g_{jj,i}) d\Omega^i,$$

$$P_{lk} = \frac{1}{16\pi G} \lim_{R \rightarrow \infty} \int_{S_R} 2(p_{ik} - \delta_{ik} p_{jj}) d\Omega^i.$$

*Remark.*

1.  $E_l$  is called the ADM energy(mass) of that end. In mathematical literature, the name mass is preferred. In physics, however, the meaning of the mass and the energy are different. The mass is a frame-independent quantity while the energy is the time component of a four vector. For the definitions in section 2, the Brown-York one is a quasilocal energy and the Liu-Yau one is a quasilocal mass. For a suitable expression of the Brown-York energy-momentum and further discussion, see [M].
2. In 1986, Bartnik showed that the definition of ADM mass is actually independent of the choice of coordinate system [B].

Although ADM mass is physically a natural candidate representing the mass of a system, the positivity of this quantity is not clear. Many physicists and mathematicians proved its positivity under additional assumptions. Finally, this problem was settled by R. Schoen and S.-T. Yau, using geometric analysis, and Witten, using spinors.

**Theorem** ([SY1, SY2]): *Let  $(M, g_{ij}, p_{ij})$  be an asymptotically flat 3-dimensional manifold satisfying  $\mu \geq |\mathbf{J}|$  in a spacetime  $N$ . Then  $E_l \geq 0$  on each end  $M_l$ . If  $E_l = 0$  for some  $l$  then  $M$  has only one end and  $M$  can be isometrically embedded into four dimensional Minkowski spacetime as a spacelike hypersurface so that  $p_{ij}$  is the second fundamental form. In particular  $M$  is topologically  $\mathbb{R}^3$ .*

**Theorem** ([W], see also [PT]). *Let  $(M, g_{ij}, p_{ij})$  be an asymptotically flat 3-dimensional manifold satisfying the dominant energy condition in a spacetime  $N$ . Then  $E_l \geq |P_l|$  on each end  $M_l$ . If  $E_l = 0$  for some  $l$  then  $M$  has only one end and  $M$  is flat along  $N$ , i.e.  $R_{\alpha\beta\gamma\delta}|_M = 0$ , where  $R_{\alpha\beta\gamma\delta}$  is the curvature tensor of  $N$*

*Remark.* For the definition of  $\mu$  and  $\mathbf{J}$ , see section 4. The condition  $\mu \geq |\mathbf{J}|$  is equivalent to the weak energy condition(See Appendix.)

It is worth some discussion on both approaches since they form the basis of the later proofs of various positive mass theorems. Schoen and Yau first prove the Riemannian case. That is, the mass of an asymptotically flat Riemannian three manifolds with nonnegative scalar curvature is nonnegative and zero only if it is isometric to  $\mathbb{R}^3$  with Euclidean metric. They reduce the general case to the Riemannian one by constructing a scalar flat Riemannian three manifold. This manifold is obtained by solving the Jang's equation and tends to the original one at the infinity. Therefore, the mass of the two manifolds are the same. Witten divides the proof into three steps. First, he derived a Weitzenböck formula for the spinor

$$\int_M |\nabla\psi|^2 + \langle\psi, \mathcal{R}\cdot\psi\rangle - |\mathcal{D}\psi|^2 = \frac{1}{2} \int_{\partial M} \langle\psi, [e^i, e^j]\cdot\nabla_j\psi\rangle e_i \lrcorner \mu \quad (1)$$

The next and usually the hardest step is to prove the existence of asymptotically constant harmonic spinors. Witten's proof in this step is not rigorous. The full proof is given in [PT, section5]. The final step is to apply the Weitzenböck formula to an asymptotically constant spinor and identify the boundary integral with  $m_{\text{ADM}}|\psi_0|^2$ .

*Remark.* Spinor is the section of the spinor bundle over  $M$ . On the end  $M_l$ , the spinor bundle is trivial. Picking one trivialization, the spinor can be viewed as a vector-valued function, so we can define what a constant spinor is. Note that the notion of constant spinor depends on coordinate.

In physics, it is also desirable to find suitable quasilocal notions of energy and momentum. We would like to define an energy-momentum tensor for a compact spacelike two-surface in spacetime. The energy-momentum vector should only depend on the first and second fundamental forms and the connection on normal bundle of the two-surface. According to Christodoulou and S.-T. Yau [CY], Melissa Liu and S.-T. Yau [LY2], the quasilocal mass should also satisfy the following properties.

- (1) It should be zero for the flat spacetime.
- (2) The quasilocal mass should be equivalent to the standard definition when evaluated on the spheres if the spacetime is spherically symmetric. In particular, for the centered spheres in the Schwarzschild spacetime, the quasilocal mass should be equivalent to the standard mass.
- (3) For an asymptotically flat slice, the quasilocal energy-momentum vector of the coordinate sphere should asymptotic to the ADM energy-momentum vector.
- (4) For an asymptotically null slice, the quasilocal energy-momentum vector of the coordinate sphere should asymptotic to the Bondi energy-momentum vector.
- (5) For an apparent horizon  $\Sigma$ , the quasilocal mass should be no less than a (universal) constant multiple of the irreducible mass which is  $\sqrt{\text{Area}(\Sigma)/16\pi}$ .



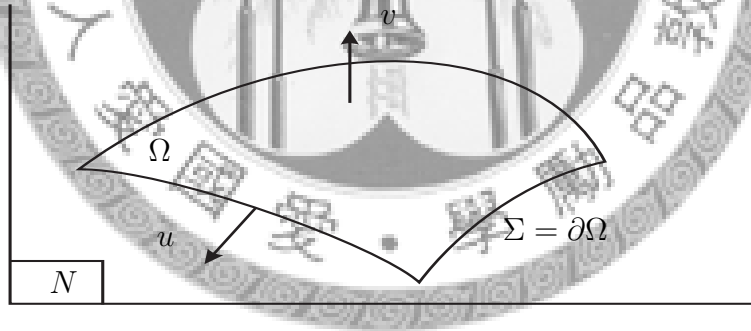
- (6) The quasilocal energy-momentum vector should be non-spacelike and the quasilocal mass should be nonnegative.

There have been many attempts to define quasilocal mass (most of them did not give the associated momentum 3-vector). Unfortunately, none of these definition could satisfy all required properties. In this note, we only discuss the Brown-York type quasilocal mass. For other definitions, the readers may consult [Sz].

The rest of the note is organized as follows. In section 2, we recall the definition of Brown-York and Liu-Yau quasilocal mass and the properties of the latter. In section 3, we describe Yu-Guang Shi and Luen-Fai Tam's approach to proving the positivity of quasilocal mass in Riemannian case. In section 4, we discuss how Liu and Yau solved the general case. In section 5, we discuss Wang and Yau's modification of Liu-Yau quasilocal mass.

## 2 The definition of Brown-York and Liu-Yau quasilocal mass

Let  $\Omega$  be a compact spacelike hypersurface with boundary in a time-oriented spacetime  $N$  with timelike future-directed unit normal  $v$ , and let  $\Sigma$  be a connected component of  $\partial\Omega$  with outward normal  $u$ . We denote the second fundamental form of  $\Omega$  in  $N$  and  $\Sigma$  in  $\Omega$  by  $p_{ij}$  and  $k_{ab}$  respectively, and  $K = \text{tr}K_{ij}$ ,  $k = \text{tr}k_{ab}$ .



*Remark.* In this note, we follow the usual convention. The Greek indices  $\alpha, \beta, \dots = 0, 1, 2, 3$ ; the Latin indices  $i, j, \dots = 1, 2, 3$ ; and  $a, b, \dots = 1, 2$ .

We need the Weyl embedding theorem:

**Theorem** (Weyl embedding theorem [Ni, Po]). *Let  $\Sigma$  be a closed surface with a Riemannian metric of positive Gauss curvature, then there exists an isometric embedding  $i : \Sigma \hookrightarrow \mathbb{R}^3$  that is unique up to Euclidean rigid motion. Furthermore,  $i(\Sigma)$  is convex.*

Suppose  $\Sigma$  has positive Gauss curvature. By Weyl embedding theorem,  $\Sigma$  can be isometrically embedded into  $\mathbb{R}^3 \subset \mathbb{R}^{3,1}$ . The second fundamental form  $(k_0)_{ab}$  of the embedded surface is positive definite and determined by the intrinsic curvature of  $\Sigma$ .

The Brown-York quasilocal mass is defined as

$$E(\Sigma, \Omega) = \frac{1}{8\pi G} \int_{\Sigma} k_0 - k.$$

If in addition the mean curvature vector  $\vec{\mathbf{H}}$  of  $\Sigma \subset N$  is spacelike, the Liu-Yau quasilocal mass is defined as

$$E(\Sigma) = \frac{1}{8\pi G} \int_{\Sigma} k_0 - |\vec{\mathbf{H}}|.$$

*Remark.* Brown and York proposed their definition in 1992 through the Hamilton-Jacobi analysis and verified its properties except positivity. In 2002, Shi and Tam proved the positivity of Brown-York mass in the time-symmetric case (see section 3). In 2003, based on Yau's work on blackholes, Liu and Yau proposed their definition out of the geometric consideration.

Liu-Yau's quasilocal mass is more intrinsic because it is independent of the three manifold  $\Sigma$  encloses. It is also a good candidate in view of the requirements mentioned in the introduction. For (1), see section 5. On the Schwarzschild spacetime,  $E(S_r) = r(1 - \sqrt{1 - \frac{2M}{r}})$  for  $r > 2M$ . Note  $E(S_{2M}) = 2M$ ,  $E(S_{\infty}) = M$ , which is consistent with (2). For (3) and (4), see [Epp].  $E(\Sigma)$  satisfies (5) by Minkowski inequality for convex surfaces. The positivity of  $E(\Sigma)$  is discussed in section 4.

### 3 The Riemannian Case

In this section, we describe Shi and Tam's proof on the positivity of quasilocal mass [ST] for the Riemannian case.

When  $\Omega$  has zero second fundamental form ( $p_{ij} = 0$ ), we say it is time-symmetric. The weak energy condition  $\mu \geq |\mathbf{J}|$  implies  $\Omega$  has nonnegative scalar curvature. The assumption that  $\Sigma$  has spacelike mean curvature vector implies  $\Sigma$  has positive mean curvature in  $\Omega$ . In this case the Brown-York and Liu-Yau quasilocal mass coincide, and the positivity of quasilocal mass reduces to a problem of Riemannian geometry.

**Theorem 1** ([ST, Theorem 4.2.]). *Let  $(\Omega^3, g)$  be a Riemannian manifold of dimension 3 with compact closure with smooth boundary and with nonnegative scalar curvature. Suppose  $\partial\Omega$  has finitely many components  $\Sigma_i$  so that each component has positive Gauss*

curvature and positive mean curvature  $H$  with respect to the outward normal. Then for each  $\Sigma_i$ ,

$$\int_{\Sigma_i} H d\sigma \leq \int_{\Sigma_i} H_0^{(i)} d\sigma.$$

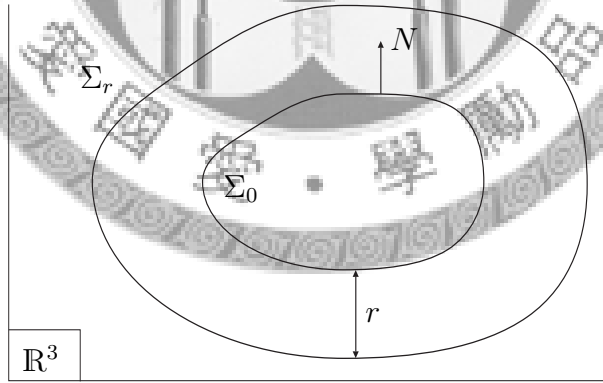
where  $H_0^{(i)}$  is the mean curvature of  $\Sigma_i$  with respect to the outward normal when it is isometrically embedded into  $\mathbb{R}^3$ . Moreover, if the equality holds for some  $\Sigma_i$ , then  $\partial\Omega$  has only one component and  $\Omega$  is isometric to a domain in  $\mathbb{R}^3$ .

*Proof(sketch)*

*Step1:* The main idea is applying Bartnik's quasi-spherical construction. Roughly speaking, we turn  $\Omega$  into a complete asymptotically flat manifold by gluing ends to  $\Omega$  and try to relate the quasilocal mass to the ADM mass of this new manifold. For simplicity, we assume  $\partial\Omega$  has only one component in the following. First we isometrically embed  $\Sigma$  into  $\mathbb{R}^3$  as a strictly convex hypersurface  $\Sigma_0$ . The position vector of the exterior  $E$  of  $\Sigma_0$  is  $\mathbf{Y} = \mathbf{X} + r\mathbf{N}$ , where  $\mathbf{X}$  is the position vector of  $\Sigma_0$  and  $\mathbf{N}$  is the unit outward normal of  $\Sigma_0$ . Let  $\Sigma_r$  be the convex hypersurface at distance  $r$  to  $\Sigma_0$ . The Euclidean space outside  $\Sigma_0$  can be represented by  $(\Sigma_0 \times [0, \infty), dr^2 + g_r)$ , where  $g_r$  is the induced metric on  $\Sigma_r$ . Next, we solve the prescribed scalar curvature equation

$$\begin{cases} 2H_0 \frac{\partial u}{\partial r} = 2u^2 \Delta_r u + (u - u^3) R^r & \text{on } \Sigma_0 \times [0, \infty) \\ u(x, 0) = u_0(x) \end{cases} \quad (2)$$

where  $u_0(x)$  is a positive smooth function on  $\Sigma_0$ , and  $H_0, R^r$  are the mean curvature and scalar curvature of  $\Sigma_r$ .



The purpose of the above construction is to deform the Euclidean metric radially to get an asymptotically flat metric while keeping the scalar curvature equal to zero. The asymptotic behavior of  $u$  also gives the ADM mass of the metric.

**Theorem** ([ST, Theorem 2.1.]). *The initial value problem (2) has a unique solution  $u$  on  $\Sigma_0 \times [0, \infty)$  such that*

- $u(z) = 1 + \frac{m_0}{\rho^{n-2}} + v$ , where  $m_0$  is a constant and  $v$  satisfies  $|v| = O(\rho^{1-n})$  and  $|\nabla_0 v| = O(\rho^{-n})$ ;
- The metric  $ds^2 = u^2 dr^2 + g_r$  is asymptotically flat with scalar curvature  $R \equiv 0$  outside  $\Sigma_0$ ;
- The ADM mass  $m_{\text{ADM}}$  of  $ds^2$  is given by

$$c(n)m_{\text{ADM}} = (n-1)\omega_{n-1}m_0 = \lim_{r \rightarrow \infty} \int_{\Sigma_r} H_0(1-u^{-1})d\sigma_r = \lim_{r \rightarrow \infty} \int_{\Sigma_r} (H_0 - H)d\sigma_r,$$

for some positive constant  $c(n)$ , where  $H_0$  and  $H$  are the mean curvatures of  $\Sigma_r$  with respect to the Euclidean metric and  $ds^2$  respectively.

*Step2:* If we view  $H(x)$  as a function on  $\Sigma_0$ , by the assumption of Theorem 1,  $\frac{H_0(x)}{H(x)}$  is positive on  $\Sigma_0$ . We solve the prescribed scalar curvature equation with initial value  $u(x, 0) = \frac{H_0(x)}{H(x)}$ , and let  $ds^2 = u^2 dr^2 + g_r$ . Let  $(M, g')$  be the Riemannian manifold obtained by gluing  $(\Omega, g)$  and  $(E, ds^2)$  along  $\Sigma \simeq \Sigma_0$ .

Note the following properties of  $g'$  :

- $g'$  is only Lipschitz near  $\partial\Omega$ .
- The mean curvature at  $\partial\Omega$  with respect to  $g'|_{N \setminus \Omega}$  and  $g'|_{\bar{\Omega}}$  coincide.
- $g'$  is asymptotically flat.
- The scalar curvature of  $N \setminus \partial\Omega$  is nonnegative (zero on  $N \setminus \bar{\Omega}$ ).

Shi and Tam are able to prove a positive mass theorem for this type of metric. The Weitzenböck formula remains the same by (ii) ([ST, Lemma 3.2.]) The existence of asymptotically constant harmonic spinor is more involved and is treated in [ST, pages 23-27]

*Step3:* The last step is to prove the monotonicity of mass expression.

**Lemma** ([ST, Lemma 4.2.]).

$$m(r) = \int_{\Sigma_r} H_0(1-u^{-1})d\sigma_r \quad \text{is nonincreasing in } r.$$

Since  $m(0) = \int_{\Sigma} H_0 - H$ ,  $m(\infty) = m_{\text{ADM}} \geq 0$ , this completes the proof of Theorem 1.

## 4 The General Case

Recall  $(\Omega, g_{ij}, p_{ij})$  refers to a compact spacelike hypersurface with boundary in a time-oriented four dimensional spacetime  $N$ , where  $g_{ij}$  and  $p_{ij}$  are the induced metric and second fundamental form of  $\Omega$ . The local mass density  $\mu$  and local current density  $J^i$  of  $\Omega$  are

$$\mu = \frac{1}{2} \left( R - \sum_{ij} p^{ij} g_{ij} + \left( \sum_i p_i^i \right)^2 \right)$$

$$J^i = \sum_j D_j (p^{ij} + \left( \sum_k p_k^k \right) g^{ij})$$

where  $R$  is the scalar curvature of  $g_{ij}$ .

**Theorem 2** ([LY2, Theorem 1]). *Suppose  $\mu$  and  $J^i$  satisfies the weak energy condition  $\mu \geq \sqrt{J^i J_i}$  and the boundary  $\partial\Omega$  has finitely many connected components  $\Sigma^1, \dots, \Sigma^l$ , each of which has positive Gaussian curvature and spacelike mean curvature vector in  $N$ . Then  $E(\Sigma^\alpha) \geq 0$  for  $\alpha = 1, \dots, l$ . Moreover, if  $E(\Sigma^\alpha) = 0$  for some  $\alpha$ , then  $N$  is flat along  $\Omega$  and  $\partial\Omega$  is connected.*

*Proof(sketch)*

*Step1:* (Construct a scalar flat three manifold)The main idea is to reduce this case into the Riemannian one. Consider the Jang's equation on  $\Omega$  with Dirichlet boundary condition:

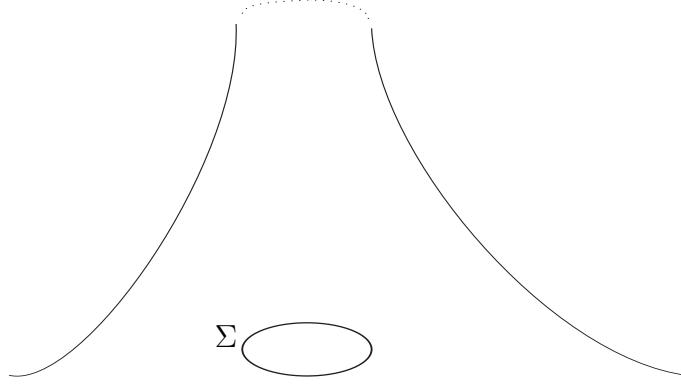
$$\begin{cases} \sum_{i,j=1}^3 (g^{ij} - \frac{f^i f^j}{1+|\nabla f|^2}) (\frac{f_{ij}}{1+|\nabla f|^2} - p_{ij}) = 0 \\ f|_{\partial\Omega} = 0 \end{cases}$$

Yau showed there exists a solution to this boundary value problem when  $(\Omega, g_{ij}, p_{ij})$  contains no apparent horizon [Y](with the main estimates in [SY2, setion 3]). When  $(\Omega, g_{ij}, p_{ij})$  has apparent horizons, the solution would blow up around the apparent horizons, but the graph of the solution in  $\Omega \times \mathbb{R}$  can be compactified to get a smooth manifold with a discontinuous metric [SY2, p.257].

Let  $\bar{g}_{ij} = g_{ij} + f_i f_j$  be a new metric that coincides with  $g_{ij}$  on  $\partial\Omega$ . The scalar curvature of  $\bar{g}$  satisfies

$$\bar{R} \geq 2|X|^2 - 2 \operatorname{div} X,$$

for some vector field  $X$  (For the explicit form of  $X$ , see [LY2, p.7]) This is enough for the existence of a scalar flat metric.



**Proposition** ([LY2, Proposition 5]). *Suppose the scalar curvature  $\bar{R}$  satisfies  $\bar{R} \geq c|X|^2 - 2 \operatorname{div} X$ , for some constant  $c > \frac{1}{2}$  and some smooth vector field  $X$  on  $\Omega$ . Then there is a unique metric  $\hat{g}_{ij}$  on  $\Omega$  such that*

1. *The metric  $\hat{g}_{ij}$  is conformal to  $\bar{g}_{ij}$ .*
2. *The scalar curvature of  $\hat{g}_{ij}$  is zero.*
3. *The metric  $\hat{g}_{ij}$  coincides with  $\bar{g}_{ij}$  on  $\partial\Omega$ .*
4. *Let  $\bar{H}$  and  $\hat{H}$  denote the mean curvature with respect to the metric  $\bar{g}$  and  $\hat{g}$  respectively, and let  $\bar{\nu}$  denote the outward unit normal of  $\partial\Omega$  in  $(\Omega, \bar{g})$ . Then*

$$\int_{\partial\Omega} \hat{H} \geq \int_{\partial\Omega} (\bar{H} - \langle X, \bar{\nu} \rangle),$$

*where the equality holds if and only if  $\bar{R} = 0$ ,  $X = 0$ , and  $\hat{g}_{ij} = \bar{g}_{ij}$ .*

*Step 2: (Glue ends to  $\Omega$ ) We have the following computation:*

**Lemma** ([LY2, Lemma 6.]).

$$\bar{H} - \langle X, \bar{\nu} \rangle \geq |\vec{\mathbf{H}}|$$

Because  $\vec{\mathbf{H}}$  is assumed to be spacelike,  $\bar{H} - \langle X, \bar{\nu} \rangle$  is positive. Liu and Yau next modified Shi and Tam's approach by solving the prescribed scalar curvature equation (2) with initial value  $h(x, 0) = \frac{H_0^\alpha}{\bar{H} - \langle X, \bar{\nu} \rangle}$  on  $E^\alpha \simeq \Sigma^\alpha \times [0, \infty)$ . Again  $m^\alpha(r) = \frac{1}{8\pi G} \int_{\Sigma_r^\alpha} (H_0 - H) d\sigma_r$  on  $(E^\alpha, g^\alpha = h^2 dr^2 + g_r)$  is nonincreasing in  $r$ . Together with the previous lemma,

$$m^\alpha(0) = \frac{1}{8\pi G} \int_{\Sigma} (H_0^\alpha - (\bar{H} - \langle X, \bar{\nu} \rangle)) d\sigma \leq \frac{1}{8\pi G} \int_{\Sigma} (H_0^\alpha - |\vec{\mathbf{H}}|) d\sigma = E(\Sigma^\alpha) \quad (3)$$

Let  $(M, \tilde{g})$  be the three manifold obtained by gluing  $(E^\alpha, g^\alpha)$  to  $(\Omega, \hat{g} = u^4 \bar{g})$ .  $\tilde{g}$  is a continuous Riemannian metric that is

1. smooth on  $M \setminus \Omega$  and  $\bar{\Omega}$ , and is Lipschitz near  $\partial\Omega$ .
2. asymptotically flat on each end  $E^\alpha$ .
3. scalar flat on  $M \setminus \partial\Omega$ .

*Step 3:* In view of (3), it suffices to prove a positive mass theorem for  $(M, \tilde{g})$ . However, two difficulties arise because of the discontinuity of mean curvature along  $\partial\Omega$ . First, a new term appears in the Weitzenböck formula:

**Lemma** ([LY2, Lemma 11.]). *Let  $U$  be an open set of  $M$ . For any spinor  $\eta \in W_0^{1,2}(U, S)$ ,  $\psi \in W_{loc}^{1,2}(U, S)$ , we have*

$$\int_U \langle \mathbf{D}\psi, \mathbf{D}\eta \rangle = \int_U \langle \nabla\psi, \nabla\eta \rangle_{\tilde{g}} + \int_{\partial\Omega \cap U} (2\bar{\nu}(u) + \frac{1}{2}\langle X, \bar{\nu}u \rangle) \langle \psi, \eta \rangle,$$

where  $u$  is the conformal factor of  $\tilde{g} = u^4\bar{g}$ .

Liu and Yau overcome these difficulties by establishing the following inequality

**Proposition** ([LY2, Proposition 10], see also [WY, Theorem 5.1] for a simpler argument). *For  $r > L$  and  $\psi \in W_{loc}^{1,2}(M, S) \cap C^\infty(M \setminus M_L, S)$ , we have*

$$2 \int_{M_r} |\mathbf{D}\psi|^2 \geq \frac{1}{10} \int_{M_r} |\nabla\psi|^2 + \frac{1}{16} \int_{\Omega} u^{-2} |du|^2 |\psi|^2 + \sum_{\alpha=1}^l \int_{S_r^\alpha} \langle \frac{H}{2}\psi - c(\nu)\check{\mathbf{D}}\psi, \psi \rangle$$

where  $\check{\mathbf{D}}$  is the Dirac operator on  $S_r^\alpha$ .

Second, the zeroth term of the Dirac operator can be discontinuous along  $\partial\Omega$ . Liu and Yau modified the argument in [PT, section 5] under weaker regularity. Indeed, the harmonic spinors lie in  $W^{1,p}(M)$  instead of  $C^\infty(M)$ , but such regularity is sufficient to prove the positive mass theorem here.

By a calculation similar to that in [PT, pages 231-232],

$$\lim_{r \rightarrow \infty} \int_{S_r^\alpha} \langle \frac{H}{2}\psi - c(\nu)\check{\mathbf{D}}\psi, \psi \rangle = -m_\infty^\alpha |\psi_0|^2$$

for a spinor asymptotic to a constant spinor  $\psi_0$ , where  $m_\infty^\alpha = \lim_{r \rightarrow \infty} m^\alpha(r)$ . This finishes the proof of Theorem 2.

## 5 A new quasilocal mass of Wang and Yau

In this final section, we present the recent work of Mu-Tao Wang and Yau on quasilocal mass [WY].

The Liu-Yau quasilocal mass does not satisfy the required property (1) in the introduction. In [MST], Murchadha, Szabados, and Tod construct some surfaces with strictly positive Liu-Yau mass lying in the lightcone of  $\mathbb{R}^{3,1}$ . In order to resolve this inconsistency, Wang and Yau proposed to take the momentum information  $p_{ij}$  into account. They take the reference to be an isometric embedding into  $\mathbb{R}^{3,1}$  instead of  $\mathbb{R}^3$ . The first task is to show the existence and uniqueness of such isometric embedding with prescribed time function.

**Theorem 3** ([WY, Theorem 3.1]). *Let  $\Sigma$  be a two-surface diffeomorphic to  $S^2$  with metric  $\sigma$ ,  $\tau$  be a function on  $\Sigma$ , and  $T_0$  be a fixed timelike vector in  $\mathbb{R}^{3,1}$ . Suppose*

$$K + (1 + |\nabla\tau|^2)^{-1} \det(\nabla^2\tau) > 0,$$

where  $K$  is the Gauss curvature of  $\Sigma$  and  $\det(\nabla^2\tau)$  is the determinant of the Hessian of  $\tau$ . Then there exists a unique spacelike embedding  $X : \Sigma \hookrightarrow \mathbb{R}^{3,1}$  with the induced metric  $\sigma$  and  $\langle X, T_0 \rangle = \tau$ .

The new quasilocal mass is defined as the difference of the *static mean curvature* between the two isometric embedding  $i : \Sigma \hookrightarrow N$ , and  $i_0 : \Sigma \hookrightarrow \mathbb{R}^{3,1}$ .

**Definition.**

- [WY, Definition 2.1] Suppose  $i : \Sigma \hookrightarrow N$  is an embedded spacelike two-surface. Given a smooth function  $\tau$  on  $\Sigma$  and a spacelike normal  $e_3$ , the static mean curvature associated with these data is defined to be

$$h(\Sigma, i, \tau, e_3) = -\sqrt{1 + |\nabla\tau|^2} \langle \vec{\mathbf{H}}, e_3 \rangle + \alpha_{e_3}(\nabla\tau)$$

where  $\vec{\mathbf{H}}$  is the mean curvature vector of  $\Sigma$  in  $N$  and  $\alpha_{e_3}(v) = \langle \nabla_v^N e_3, e_4 \rangle$  is the connection form of the normal bundle of  $\Sigma$  in  $N$  determined by  $e_3$  and the future-directed timelike unit normal  $e_4$  orthogonal to  $e_3$ .

- [WY, Definition 2.2] Given an isometric embedding  $i : \Sigma \hookrightarrow N$  with spacelike mean curvature vector  $\vec{\mathbf{H}}$ . Denote

$$\mathfrak{H}(\Sigma, i, \tau) = \int_{\Sigma} h(\Sigma, i, \tau, \bar{e}_3) dv_{\Sigma},$$

where  $h(\Sigma, i, \tau, \bar{e}_3) = \min_{e_3} \{h(\Sigma, i, \tau, e_3)\}$ .



3. [WY, Definition 5.2] Given a spacelike embedding  $i : \Sigma \hookrightarrow N$ . Suppose the set of admissible functions is non-empty (See [WY, Definition 5.1]). The quasilocal mass is defined to be the infimum of

$$\mathfrak{H}(\Sigma, i_0, \tau) - \mathfrak{H}(\Sigma, i, \tau)$$

among all admissible  $\tau$ , where  $i_0$  is the unique spacelike isometric embedding of  $\Sigma$  into  $\mathbb{R}^{3,1}$  associated with  $\tau$  given by Theorem 3.

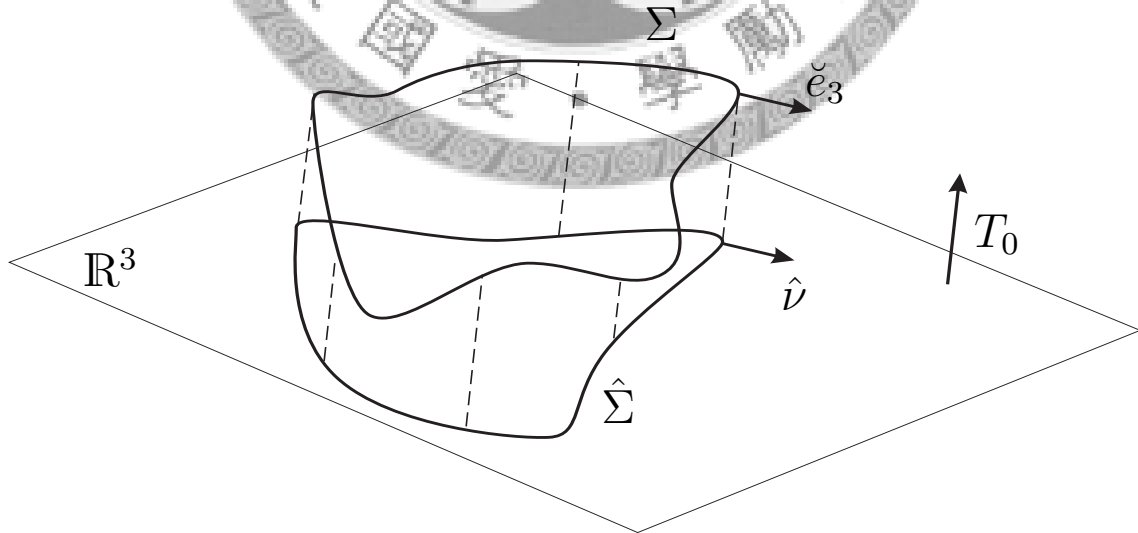
For a two-surface  $\Sigma \subset \mathbb{R}^{3,1}$ ,  $i = i_0$ . If the projection of  $\Sigma$  along some time direction is a convex surface, then  $\Sigma$  has zero quasilocal mass. This case covers the examples of Murchadha, Szabados, and Tod.

We briefly mention the idea of proving the positivity of this new quasilocal mass. For an embedded two-surface  $\Sigma \subset \mathbb{R}^{3,1}$ , we denote by  $\hat{\Sigma}$  its projection onto  $\mathbb{R}^3$ . The most important observation of Wang and Yau is to identify the two terms  $\int_{\Sigma} k_0$  and  $\int_{\Sigma} H - \langle X, \nu \rangle$  appearing in Liu and Yau's proof to the integral of some static mean curvature.

**Theorem 4** ([WY, Proposition 3.1, 3.2]).

$$\int_{\hat{\Sigma}} \hat{k} = \int_{\Sigma} h(\Sigma, i_0, \tau, \check{e}_3) dv_{\Sigma}.$$

where  $\hat{k}$  is the mean curvature of  $\hat{\Sigma}$  in  $\mathbb{R}^3$  with respect to the outward normal  $\hat{\nu}$ , and  $\check{e}_3$  is obtained by parallel translating  $\hat{\nu}$  along  $T_0$ . Furthermore, when the mean curvature of  $\Sigma$  in  $\mathbb{R}^{3,1}$  is spacelike,  $\int_{\hat{\Sigma}} \hat{k} dv_{\hat{\Sigma}} = \mathfrak{H}(\Sigma, i_0, \tau)$ .



**Theorem 5** ([WY, Theorem 4.1]). *Let  $i : \Sigma \hookrightarrow N$  be a spacelike embedding. Given any smooth function  $\tau$  on  $\Sigma$  and any spacelike hypersurface  $\Omega$  with  $\partial\Omega = \Sigma$ . Suppose the Dirichlet problem of the Jang's equation over  $\Omega$  subject to the boundary condition that  $f = \tau$  on  $\Sigma$  is solvable. Then there exists a spacelike unit normal  $e'_3$  along  $\Sigma$  in  $N$  such that the expression  $\tilde{k} - \langle \tilde{\nabla}_{\tilde{e}_4} \tilde{e}_4, \tilde{e}_3 \rangle + P(\tilde{e}_4, \tilde{e}_3)$  (this is the familiar term  $\hat{H} - \langle X, \nu \rangle$  in Liu and Yau's paper) at  $\tilde{q} \in \tilde{\Sigma} \subset \Omega \times \mathbb{R}$  is equal to*

$$(1 + |\nabla\tau|^2)^{-1/2} h(\Sigma, i, \tau, e'_3) \text{ at } q \in \Sigma,$$

where  $\tilde{q} = (q, \tau(q)) \in \tilde{\Sigma}$ .

Combining these two theorems and the result of Liu and Yau,

$$\begin{aligned} \mathfrak{H}(\Sigma, i_0, \tau) &= \int_{\tilde{\Sigma}} \tilde{k} \\ &\geq \int_{\tilde{\Sigma}} \tilde{k} - \langle \tilde{\nabla}_{\tilde{e}_4} \tilde{e}_4, \tilde{e}_3 \rangle + P(\tilde{e}_4, \tilde{e}_3) \\ &= \int_{\Sigma} h(\Sigma, i, \tau, e'_3) dv_{\Sigma} \\ &\geq \mathfrak{H}(\Sigma, i, \tau) \end{aligned}$$

Suppose the two-surface  $\Sigma$  bounds a spacelike hypersurface in  $N$ , and has positive Gauss curvature and spacelike mean curvature vector. Then the assumptions of Theorem 3 and Theorem 5 are satisfied (For details, see [WY, Theorem 4.2].) We can conclude that

**Theorem** ([WY, Corollary 5.3]). *Under the assumption of Theorem 2, the new quasilo-cal mass is nonnegative.*

## Appendix

It is a well-known fact that the weak energy condition is equivalent to  $\mu \geq |\mathbf{J}|$ . We just write it down in this appendix for completion.

We first fix the notation. Let  $N$  be a 4-manifold with a Lorentzian metric of signature  $(-+++)$ . For a point  $x \in N$ , and an orthonormal frame  $\{e_{\alpha} \mid \alpha = 0, \dots, 3\}$  near  $x$ , we denote the curvature tensor by

$$\begin{aligned} R(e_{\alpha}, e_{\beta})e_{\gamma} &= \nabla_{e_{\alpha}} \nabla_{e_{\beta}} e_{\gamma} - \nabla_{e_{\beta}} \nabla_{e_{\alpha}} e_{\gamma} - \nabla_{[e_{\alpha}, e_{\beta}]} e_{\gamma} = R_{\alpha\beta\gamma}{}^{\delta} e_{\delta}, \\ R_{\alpha\beta\gamma\delta} &= g_{\sigma\delta} R_{\alpha\beta\gamma}{}^{\sigma}. \end{aligned}$$

Note  $R_{\alpha\beta\gamma 0} = -R_{\alpha\beta\gamma}{}^0$ . The Ricci curvature and scalar curvature are

$$\begin{aligned} R_{\alpha\beta} &= R_{\gamma\alpha\beta}{}^\gamma = -R_{0\alpha\beta 0} + R_{1\alpha\beta 1} + R_{2\alpha\beta 2} + R_{3\alpha\beta 3}, \\ R &= g^{\alpha\beta} R_{\alpha\beta} = -R_{00} + R_{11} + R_{22} + R_{33}. \end{aligned}$$

Suppose we have a spacelike hypersurface  $M \subset N$  with the induced metric. We denote the connection and curvature of  $N$  and  $M$  by  $\bar{D}, \bar{R}_{\alpha\beta\gamma\delta}$  and  $D, R_{ijkl}$  respectively. Let  $e_0$  be the timelike unit normal of  $M$ . In the neighborhood of a fix point  $x \in M$ , we choose a normal frame  $\{e_i \mid i = 1, \dots, 3\}$  at  $x$  diagonalizing the second fundamental form, that is, for any  $i, j$ ,  $D_{e_i}e_j(x) = 0$ ,  $p_{ij}(x) = p(e_i, e_j) = k_i\delta_{ij}$ .

We compute the Gauss and Codazzi equations for hypersurfaces:

$$\begin{aligned} \langle R(X, Y)Z, W \rangle &= \langle \bar{R}(X, Y)Z, W \rangle - p(X, W)p(Y, Z) + p(X, Z)p(Y, W) \\ -\langle \bar{R}(X, Y)Z, e_0 \rangle &= D_X p(Y, Z) - D_Y p(X, Z). \end{aligned}$$

where  $X, Y, Z, W$  are tangent vectors of  $M$ .

*Proof.* For Gauss equation,

$$\begin{aligned} \langle D_X D_Y Z, W \rangle &= \langle \bar{D}_X \bar{D}_Y Z, W \rangle \\ &= \langle \bar{D}_X (\bar{D}_Y Z + \langle \bar{D}_Y Z, e_0 \rangle e_0), W \rangle, \quad \text{since } \langle e_0, e_0 \rangle = -1 \\ &= \langle \bar{D}_X \bar{D}_Y Z, W \rangle + \langle \bar{D}_Y Z, e_0 \rangle \langle \bar{D}_X e_0, W \rangle \\ &= \langle \bar{D}_X \bar{D}_Y Z, W \rangle - p(Y, Z)p(X, W) \end{aligned}$$

The computation of  $\langle D_X D_Y Z, W \rangle$ ,  $\langle D_{[X, Y]} Z, W \rangle$  is similar.

For Codazzi equation,

$$\begin{aligned} D_X p(Y, Z) &= X p(Y, Z) - p(\bar{D}_X Y, Z) - p(Y, \bar{D}_X Z) \\ &= -X \langle \bar{D}_Y Z, e_0 \rangle - \langle \bar{D}_Z e_0, D_X Y \rangle - \langle \bar{D}_Y e_0, D_X Z \rangle \\ &= -\langle \bar{D}_X \bar{D}_Y Z, e_0 \rangle - \langle \bar{D}_Y Z, \bar{D}_X e_0 \rangle - \langle \bar{D}_Z e_0, D_X Y \rangle - \langle \bar{D}_Y e_0, D_X Z \rangle \\ -D_Y p(X, Z) &= \langle \bar{D}_Y \bar{D}_X Z, e_0 \rangle + \langle \bar{D}_X Z, \bar{D}_Y e_0 \rangle + \langle \bar{D}_Z e_0, D_Y X \rangle + \langle \bar{D}_X e_0, D_Y Z \rangle \end{aligned}$$

Canceling the second and fourth terms and combining the third term,

$$\langle \bar{D}_Z e_0, D_Y X - D_X Y \rangle = -\langle \bar{D}_Z e_0, [X, Y] \rangle = \langle e_0, \bar{D}_Z [X, Y] \rangle = \langle e_0, \bar{D}_{[X, Y]} Z \rangle,$$

we get the desired result.

To verify our claim, it is sufficient to show  $\mu = T_{00}$ ,  $J_i = T_{0i}$ , up to a constant. From

Einstein equation  $T_{\alpha\beta} = R_{\alpha\beta} - \frac{1}{2}Rg_{\alpha\beta}$  (We omit the constant  $8\pi G$ ),

$$\begin{aligned}
T_{00} &= \bar{R}_{00} + \frac{1}{2}\bar{R} \\
&= \frac{1}{2}(\bar{R}_{00} + \bar{R}_{11} + \bar{R}_{22} + \bar{R}_{33}) \\
&= \frac{1}{2}(2\bar{R}_{1221} + 2\bar{R}_{1331} + 2\bar{R}_{2332}) \\
&= \frac{1}{2}(2R_{1221} + 2p_{11}p_{22} + 2R_{1331} + 2p_{11}p_{33} + 2R_{2332} + 2p_{22}p_{33}) \\
&\qquad\qquad\qquad \text{by Gauss equation and } p_{ij} = k_i\delta_{ij} \\
&= \frac{1}{2}(R - (p_{11}^2 + p_{22}^2 + p_{33}^2) + (p_{11} + p_{22} + p_{33})^2) \\
J_i &= \sum_{j=1}^3 D_j p_{ij} - e_i \left( \sum_{k=1}^3 p_{kk} \right) \\
&= \sum_{j=1}^3 D_i p_{jj} - \bar{R}_{ijj0} - e_i \left( \sum_{k=1}^3 p_{kk} \right) \qquad \text{by Codazzi equation} \\
&= \bar{R}_{i0}
\end{aligned}$$

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